

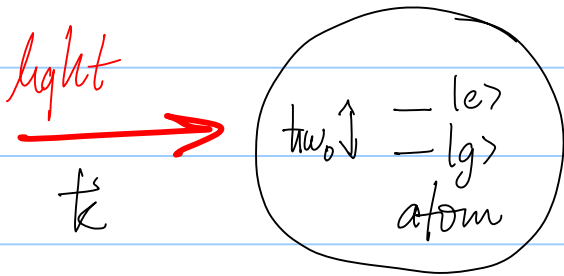
Semiclassical - quantum atom, classical light

PH/5514

Note Title

3/2/2005

Try to improve our model!



There is much, much physics here!

$$\vec{E} = E_1 \hat{e} \cos(kz - \omega t + \varphi) = \frac{1}{2} E_1 \hat{e} \left[e^{i(\vec{k} \cdot \vec{r} - \omega t + \varphi)} + e^{-i(\vec{k} \cdot \vec{r} - \omega t + \varphi)} \right]$$

What is the interaction?

$$H = H_0 + H_I$$

$$H_I = -\vec{p} \cdot \vec{E}$$

electric dipole!

$$= e \frac{E_1}{2} \vec{r} \cdot \hat{e} \left[e^{i(\vec{k} \cdot \vec{r} - \omega t + \varphi)} + e^{-i(\vec{k} \cdot \vec{r} - \omega t + \varphi)} \right]$$

Let's take our H_0 as the Hamiltonian for an atom in free space, w/ eigenkets $|g\rangle$ and $|e\rangle$ and $E_g = 0$, $E_e = \hbar\omega_0$

(we'll make life more complicated shortly)

Ah, right now: $e^{i\vec{k}\cdot\vec{r}} = 1 + i\vec{k}\cdot\vec{r} + \dots \approx 1$ since the electron is very localized compared with the wavelength of light

(we'll relax this shortly too!)

So, in the interaction picture:

$$\begin{aligned}
 H_I &= e^{iH_0 t/\hbar} H_I e^{-iH_0 t/\hbar} \\
 &= \frac{e}{2} e^{iH_0 t/\hbar} \left[e^{-i(\omega t - \phi)} + e^{i(\omega t - \phi)} \right] \vec{r} \cdot \hat{\epsilon} e^{-iH_0 t/\hbar}
 \end{aligned}$$

How do we evaluate $\vec{r} \cdot \hat{\Sigma}$?

• for a complicated atom, this is tricky — more soon

• we can learn a lot by considering a hydrogen atom with a spinless electron and nucleus

• we'll need matrix elements such as

$$\langle e | \vec{r} \cdot \hat{\Sigma} | g \rangle$$

angular integral
(selection rules)

Typically we write $\langle e | \vec{r} \cdot \hat{\Sigma} | g \rangle = D_{12} I_{\text{ang}}$

dipole matrix
element

so, for our example:

$$D_{12} = \int_0^\infty R_{n_2, l_2} r R_{n_1, l_1} r^2 dr$$

$$I_{\text{ang}} = \int d\Omega Y_{l_2 m_2}^* \hat{r} \cdot \hat{\Sigma} Y_{l_1 m_1}$$

let's rewrite $\vec{r} \cdot \hat{\zeta}$

(this is all consistent with spherical tensor notes)

we need:

$$\textcircled{1} \vec{r} = \sin\theta \cos\varphi \hat{x} + \sin\theta \sin\varphi \hat{y} + \cos\theta \hat{z}$$

$$\textcircled{2} \sin\theta \cos\varphi = \sqrt{\frac{2}{3}} (\gamma_{1,-1} - \gamma_{1,1})$$

$$\sin\theta \sin\varphi = \sqrt{\frac{2}{3}} (\gamma_{1,-1} + \gamma_{1,1})$$

$$\cos\theta = \sqrt{\frac{4}{3}} \gamma_{1,0}$$

$$\textcircled{3} \hat{\zeta} = \underbrace{\sum_{-} \left(\frac{\hat{x} - i\hat{y}}{\sqrt{2}} \right)}_{\sigma_{-}} + \underbrace{\sum_0 \hat{z}}_{\pi} + \underbrace{\sum_{+} \left(-\frac{\hat{x} + i\hat{y}}{\sqrt{2}} \right)}_{\sigma_{+}}$$

atomic
conventions:

taking this dot product in cartesian coordinates
(i.e., keeping things simple)

$$\vec{r} \cdot \hat{\zeta} = \sum_{-} \gamma_{1,-1} + \sum_0 \gamma_{1,0} + \sum_{+} \gamma_{1,1}$$

for $\mathcal{E}_0$, " π transitions"

$\hat{\mathcal{E}}$ along $\hat{z}$

$$I_{\text{ang}} = \int d\Omega Y_{\ell_2 m_2}^* \cos \theta Y_{\ell_1 m_1}$$

See Foot or many other sources -

$$I_{\text{ang}} \neq 0 \text{ iff } m_1 - m_2 = 0, \quad \boxed{\Delta m = 0}$$

for $\mathcal{E}_{\pm}$, " σ transitions"

$\hat{\mathcal{E}} \perp \hat{z}$

$$I_{\text{ang}} = \int d\Omega Y_{\ell_2 m_2}^* \sin \theta e^{\pm i\phi} Y_{\ell_1 m_1}$$

for $I_{\text{ang}} \neq 0$, require $m_1 - m_2 \pm 1 = 0$

$$\boxed{\Delta m = \pm 1}$$

Using various recursion relations, these integrals can be evaluated (Griffiths p 753)

There is an additional selection rule that you find if you do these integrals — the "parity selection rule"

This can be motivated physically —

H commutes with the parity operator,
so I_{ang} should not change under parity:

$$\hat{P} Y_{lm} = (-1)^l Y_{lm}$$

← just a property of spherical harmonics

under parity, $I_{ang} \rightarrow I_{ang} (-1)^{l_2 + l_1 + 1}$,

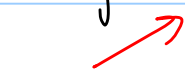
which requires

$$l_2 + l_1 + 1 = 2n \quad n = 0, 1, 2, \dots$$

The spherical harmonics Y_{lm} have angular momentum $l \geq 1$, and so we require $l_2 + l_1 = \pm 1$, per usual angular momentum addition rules.

All of these selection rules are straightforward if you evaluate I_{any} directly (see Arfken p 753):

$$I_{\text{any}} = \delta_{m_2, m_1} \left[\delta_{l_2, l_1+1} \sqrt{\frac{(l_1-m_1+1)(l_1+m_1+1)}{(2l_1+1)(2l_1+3)}} + \delta_{l_2, l_1-1} \sqrt{\frac{(l_1-m_1)(l_1+m_1)}{(2l_1-1)(2l_1+1)}} \right] \epsilon_0$$

π 

$$\delta_{m_2, m_1 \pm 1} \left[\mp \delta_{l_2, l_1+1} \sqrt{\frac{(l_1 \mp m_1+1)(l_1 \mp m_1+2)}{(2l_1+1)(2l_1+3)}} \pm \delta_{l_2, l_1-1} \sqrt{\frac{(l_1 \mp m_1)(l_1 \mp m_1-1)}{(2l_1-1)(2l_1+1)}} \right] \epsilon_{\pm}$$

$\sigma_{\pm}$ 

- keep things simple for now with the time-dependent problem -

say we only have E_+ , and therefore write

$$\hat{E} \cdot \vec{r} = (e|r_+|g\rangle|e\rangle\langle g| + \langle g|r_+|e\rangle|g\rangle\langle e|)$$

or:

$$E_{r_+} = \langle r \rangle \{E_+|e\rangle\langle g| + E_+^*|g\rangle\langle e|\}$$

$$\text{where } \langle r \rangle = \langle e|r_+|g\rangle$$

$$\text{Then: } H_I' = \langle r \rangle \frac{e}{2} E_+ e^{iH_0 t / \hbar} \left(E_- |e\rangle \langle g| + E_+^* |g\rangle \langle e| \right) e^{-iH_0 t / \hbar} \\ \cdot \left[e^{-i(\omega t - \varphi)} + e^{i(\omega t - \varphi)} \right]$$

$$= \langle r \rangle \frac{e}{2} E_+ \left[e^{i\omega_0 t} E_- |e\rangle \langle g| + e^{-i\omega_0 t} E_+^* |g\rangle \langle e| \right] \\ \left[e^{-i(\omega t - \varphi)} + e^{i(\omega t - \varphi)} \right]$$

look familiar? it's a coupled two-level system!

$$H_I' = \frac{e}{2} E_+ \langle r_+ \rangle \left[e^{i\omega_0 t} \sigma_+ E_- + e^{-i\omega_0 t} \sigma_- E_+^* \right] \\ \left[e^{-i(\omega t - \varphi)} + e^{i(\omega t - \varphi)} \right]$$

We can make the RWA:

$$H_I' \approx \frac{e}{2} E_+ \langle r_+ \rangle \left[e^{-i(\delta t - \varphi)} \sigma_+ E_- + e^{i(\delta t - \varphi)} \sigma_- E_+^* \right]$$

where $\delta = \omega - \omega_0$

and $\sigma_+ = |e\rangle \langle g|$ and $\sigma_- = |g\rangle \langle e|$

write $|\psi\rangle = c_g |g\rangle + c_e |e\rangle$

Schrödinger's equation:

$$i\hbar \frac{\partial}{\partial t} |\psi\rangle = H_I |\psi\rangle$$

*we're using the Interaction picture,
but we're not using time-dependent
PT—this is an exact solution!

$$\begin{aligned} i\hbar \dot{c}_g &= \hbar \frac{\Omega}{2} e^{i(\delta t + \phi)} c_e \\ i\hbar \dot{c}_e &= \hbar \frac{\Omega}{2} e^{-i(\delta t + \phi)} c_g \end{aligned}$$

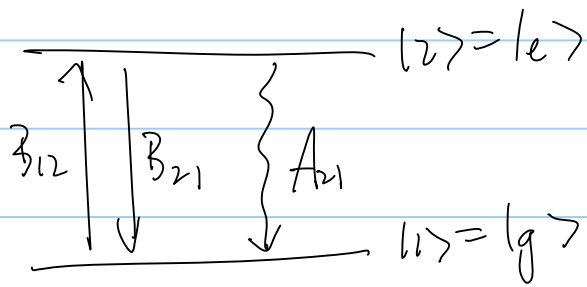
where $\hbar\Omega = e E_1 \langle r \rangle$
(this is Foot's, Metcalf's, etc.
convention)

We already know what happens! Rabi oscillations.

(we can drive rotations, etc)

Let's add in spontaneous emission ad-hoc, even though we can derive the optical Bloch equations that follow using quantum treatment of light field (we'll do that later).

Interestingly, Einstein's A & B coefficients argument is another way to get to the OBEs with minimal assumptions about the quantum nature of light.



$$A_{21} = \Gamma = 1/\tau \quad \text{lifetime of } |e\rangle$$

Einstein's argument: $A_{21} = \frac{\hbar \omega^3}{\pi^2 c^3} B_{21}$ $B_{12} = B_{21}$

let's convert to Foot's notation:

$$\Omega = \frac{e}{\hbar} \langle g | \vec{r} \cdot \vec{E}_0 | e \rangle = \frac{e}{\hbar} \langle 1 | \vec{r} \cdot \vec{E}_0 | 2 \rangle$$

$$B_{12} = \frac{\pi e^2 |D_{12}|^2}{3 \epsilon_0 \hbar^2}$$

$$\vec{D}_{12} = \langle 1 | \vec{r} | 2 \rangle \approx \text{few } a_0 \text{ for alkalis}$$

this assumes all possible polarizations

so $A_{12} = \frac{g_1}{g_2} \frac{e^2 \omega^3}{3 \pi \epsilon_0 \hbar c^3} |D_{12}|^2$ after some math

Amazing! Spontaneous emission rate depends only on ω^3 and $|D_{12}|^2$ — note ω dependence different than classical result!

We put this by hand into the OBEs:

$$\dot{c}_e = \dots - \Gamma/2 c_e$$

Note: Einstein's argument really about $|\dot{c}_e|^2$, but, if we put this into ρ_{ee} then we get an unphysical result (the coherences do not decay)

This is like a T_1 and T_2 process combined...

Foot uses a standard 'rotating frame' (it's not) operator transformation:

$$\tilde{c}_g = c_g e^{-i\delta t/2} \quad \tilde{c}_e = e^{i\delta t/2} c_e$$

$$\rightarrow \tilde{\rho}_{gg} = \rho_{gg} \quad \tilde{\rho}_{ee} = \rho_{ee} \quad \tilde{\rho}_{ge} = e^{-i\delta t} \rho_{ge} \quad \tilde{\rho}_{eg} = e^{i\delta t} \rho_{eg}$$

then:

$$\left\{ \begin{array}{l} \dot{u} = \delta v - \Gamma/2 u \\ \dot{v} = -\delta u + \Omega w - \Gamma/2 v \\ \dot{w} = -\Omega v - \Gamma(w-1) \end{array} \right. \quad \text{or} \quad \left\{ \begin{array}{l} \dot{\rho}_{gg} = \Gamma \rho_{ee} + i/2 (\Omega^* \tilde{\rho}_{eg} - \Omega \tilde{\rho}_{ge}) \\ \dot{\rho}_{ee} = -\Gamma \rho_{ee} + i/2 (\Omega \tilde{\rho}_{ge} - \Omega^* \tilde{\rho}_{eg}) \\ \dot{\tilde{\rho}}_{ge} = -(\Gamma/2 + i\delta) \tilde{\rho}_{ge} + i/2 \Omega^* (\rho_{ee} - \rho_{gg}) \\ \dot{\tilde{\rho}}_{eg} = -(\Gamma/2 - i\delta) \tilde{\rho}_{eg} + i/2 \Omega (\rho_{gg} - \rho_{ee}) \end{array} \right.$$

let's look at the long-time / equilibrium behavior:
(set all time derivatives to zero and solve)

$$\begin{pmatrix} u \\ v \\ w \end{pmatrix} = \frac{1}{\delta^2 + \Omega^2/2 + \Gamma^2/4} \begin{pmatrix} \Omega\delta \\ \Omega\Gamma/2 \\ \delta^2 + \Gamma^2/4 \end{pmatrix}$$

$$\rho_{22} = \frac{1-w}{2}$$

define: $s = \frac{2\Omega^2}{\Gamma^2} = \frac{I}{I_{\text{sat}}}$ $I_{\text{sat}} = \frac{hc\pi}{3\lambda^3\tau}$

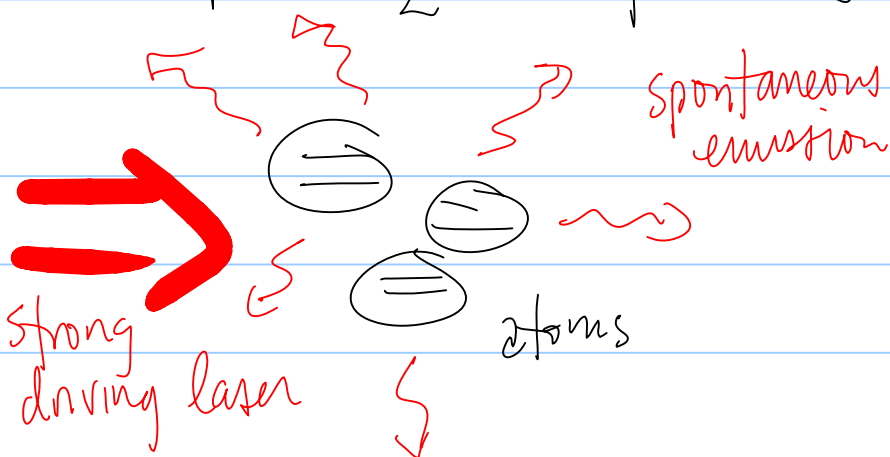
then: $\rho_{22} = \frac{1}{2} \frac{s}{1+s+4\delta^2/\Gamma^2}$

(using $I = \epsilon_0 c |\vec{E}_0|^2/2$)

rate of spontaneous emission = $\Gamma\rho_{22} = R$

first, notice that in the $s \gg 1$ limit and $\delta = 0$,

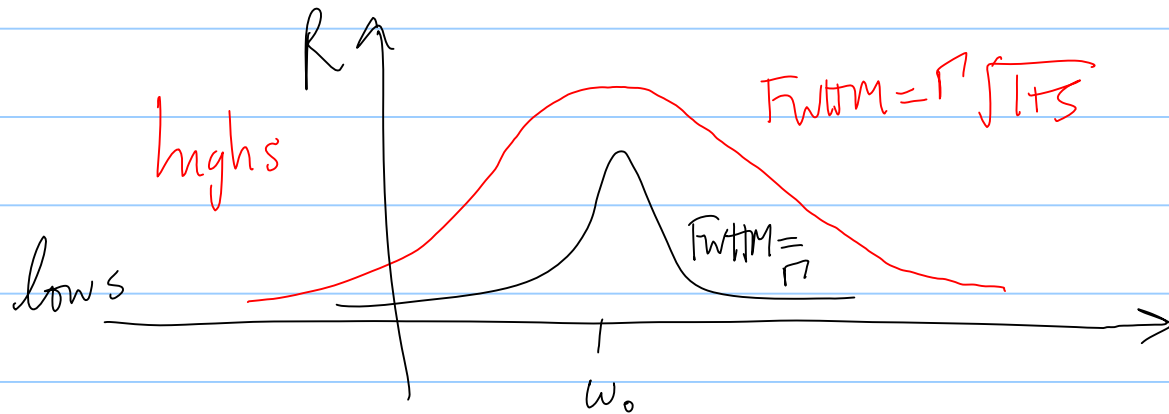
$$R \rightarrow \Gamma/2 \text{ and } \rho_{22} \rightarrow 1/2$$



the atoms spend, at most, half of their time in the excited state

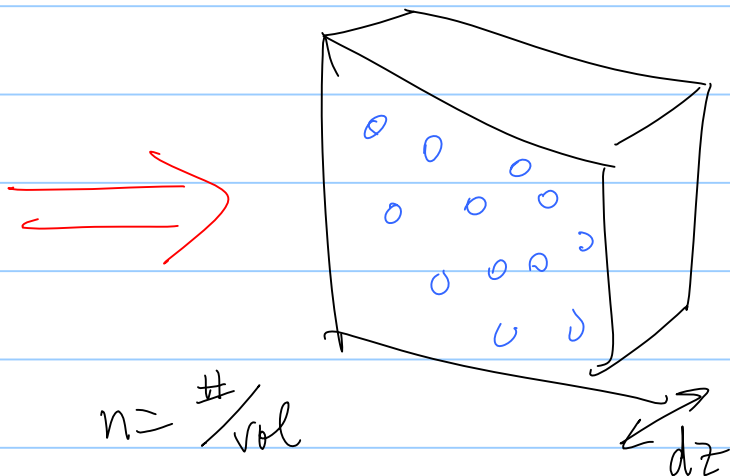
Power broadening

$$R = \frac{\Gamma}{2} s \frac{\frac{\Gamma^2}{4}}{\frac{\Gamma^2}{4}(1+s) + (\omega - \omega_0)^2}$$



physics: near resonance, very saturated — fluorescence reduced; away from resonance — not much change

What about absorption?



$$\# \text{ atoms/area} = n dz$$

$n dz =$ fraction of surface covered by atoms

$$\frac{dI}{I} = -n\sigma dz \quad \text{fraction of light absorbed}$$

$$\frac{dI}{dz} = -n\sigma I \quad \text{this defines the cross section } \sigma$$

using steady state conservation of energy:

$$\underbrace{(n_1 - n_2)\sigma I}_{\text{net rate of energy in}} = \underbrace{n_2 \Gamma \hbar \omega}_{\text{rate of energy out}}$$

n_1 : density of atoms in state 1

$$\text{using } w = \frac{n_2 - n_1}{n} \quad \rho_{22} = \frac{n_2}{n} \quad n = n_1 + n_2$$

$$\sigma = \frac{3\lambda^2}{2\pi} \frac{1}{1 + 4\delta^2/\Gamma^2}$$

$$\text{on resonance, } \sigma = \frac{3\lambda^2}{2\pi} \simeq \frac{\lambda^2}{2}$$

that's a big cross-section for something so small!

Note! There is some averaging here to be careful about.

Our definition of B_{12} has a factor of $1/3$ for average over polarization.

Everyone handles this differently. The right way to deal with this is to solve the OBEs for a real, multilevel atom, using the exact dipole matrix elements.

A reasonable approach from Foot:

$$G = \underbrace{\frac{g_2}{g_1} \frac{\pi^2 c^2}{\omega_0^2} A_{21}}_{\text{degeneracy factors from } A_{21}} \underbrace{\frac{1}{2\pi} \frac{\Gamma}{\delta^2 + \Gamma^2/4}}_{\text{lineshape function}}$$

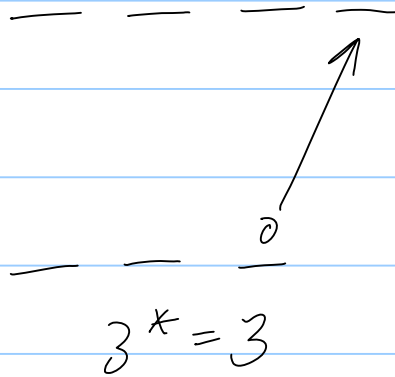
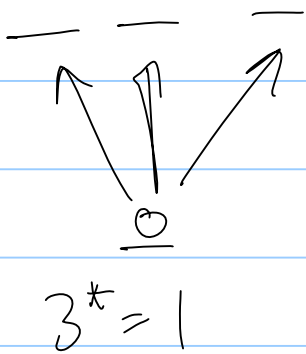
→ good for low excited state populations

A lot of people put averages into I_{sat} :

$$I_{\text{sat}} = \frac{hc\pi}{3^* \lambda^3 \tau} \quad 3^* = 1 \rightarrow 3$$

unpolarized light or atoms: $3^* = 1$

polarized light and atoms: $3^* = 3$



for alkali atoms, $I_{\text{sat}} \sim 5-6 \text{ mW/cm}^2$

So, what is the right way to calculate dipole matrix elements?

We want: $\langle F'm_F' | \vec{r} | Fm_F \rangle$

This is not a fn matrix element to calculate!

procedure:

- ① expand to $|m_J, m_I\rangle$ basis
- ② then expand to $|m_L, m_S, m_I\rangle$ basis
- ③ then calculate the matrix element

result:

$$\langle F'm_F' | r_q | Fm_F \rangle = (-1)^{L'+S+J+J'+I-m_F'} \langle \alpha' L' || r || \alpha L \rangle \cdot \begin{Bmatrix} L' & J' & S \\ J & L & 1 \end{Bmatrix} \begin{Bmatrix} J' & F' & I \\ F & J & 1 \end{Bmatrix} \cdot \begin{pmatrix} F & 1 & F' \\ m_F & q & -m_F' \end{pmatrix}$$

reduced matrix element

6-j symbols

↑ selection rules come from this 3-j symbol

The reduced matrix element is:

$$R = \int_0^\infty r^2 dr R_{n'd'} r R_{ne}$$

and must, in general, be determined experimentally
(e.g., from σ)

Mathematica knows 3-j and 6-j symbols:

<http://mathworld.wolfram.com/Wigner6j-Symbol.html>

<http://mathworld.wolfram.com/Wigner3j-Symbol.html>

A standard reference on this is Elementary Theory of Angular Momentum by Rose:

<http://www.amazon.com/Elementary-Theory-Angular-Momentum-M-E/dp/0486684806>

selection rules: * the polarization in the atomic basis matters (relative to a quantization magnetic field)

$$m_F' = q + m_F$$

$$\Delta l = \pm 1 \quad (\text{remember "parity rule"})$$

$$\Delta F = \pm 1, 0 \quad ; \quad \Delta F = 0 \text{ and } \Delta m_F = 0 \text{ not allowed}$$

No one in their right mind calculates this directly!

Talk to your favourite atomic physicist to get a chart!

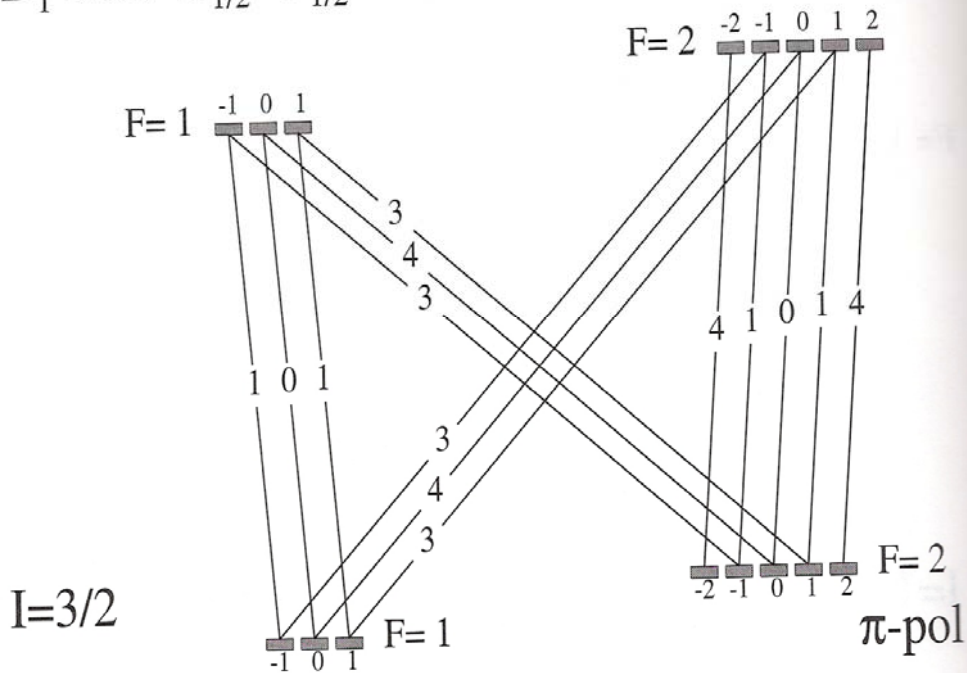
Notes on charts:

- everyone has their own conventions
- most people leave out signs - do this at your own peril!

Lesson - be sure that you know what someone means!

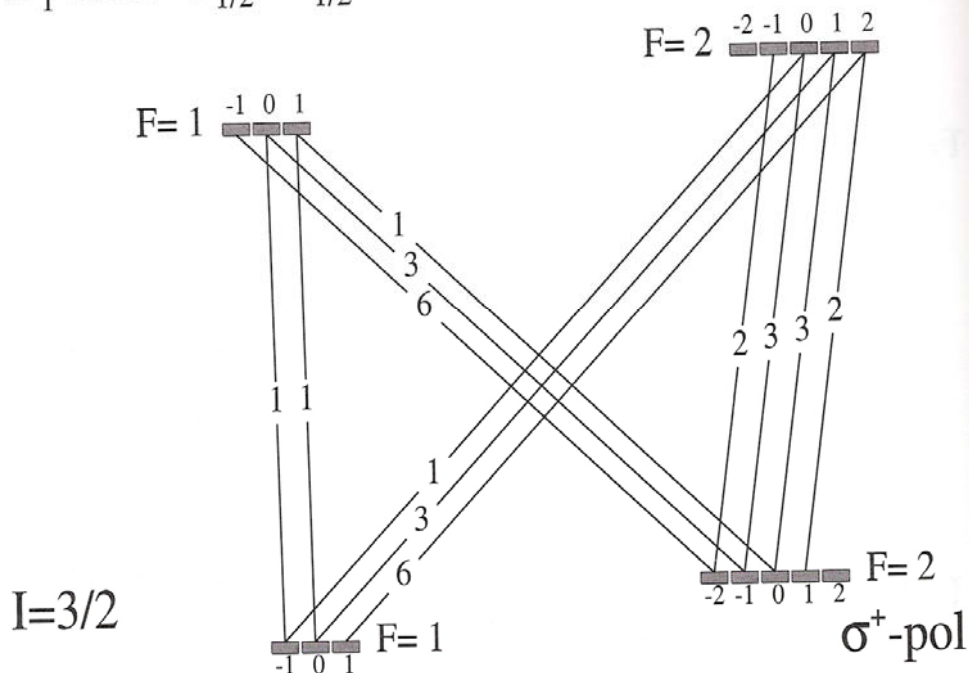
Metcalf : laser cooling and trapping

D₁-line: $^2S_{1/2} - ^2P_{1/2}$

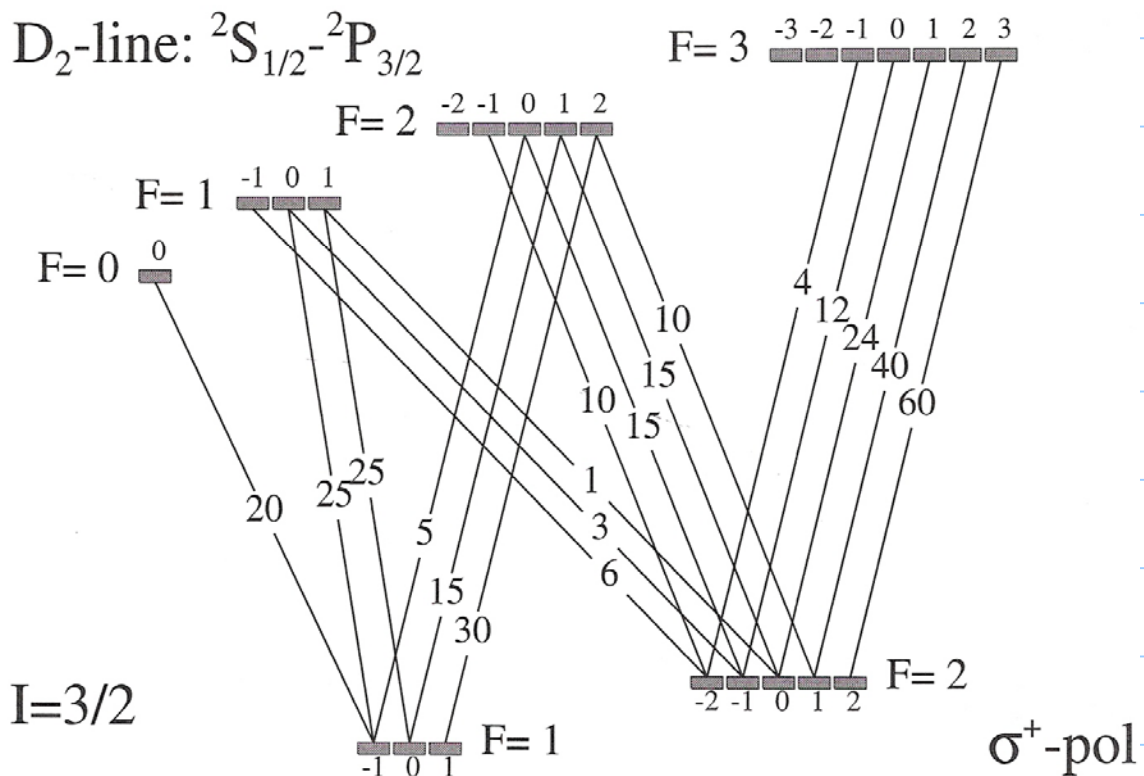
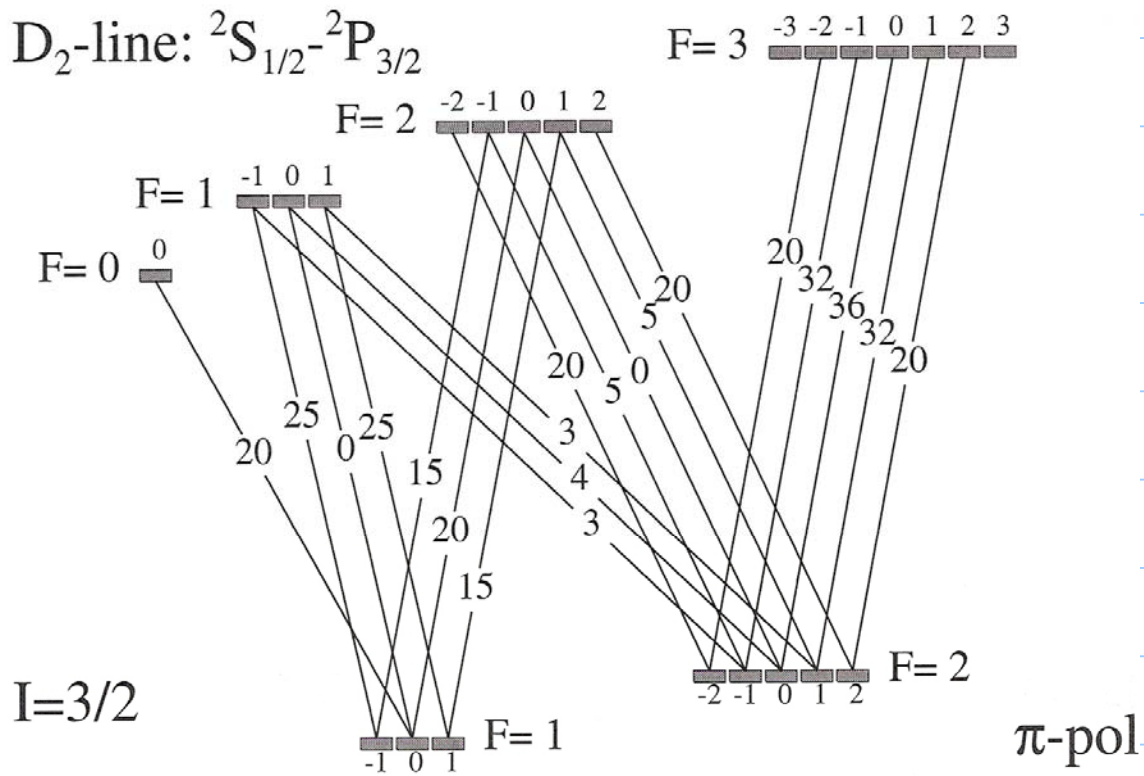


take the
square
root -
watch
normalization

D₁-line: $^2S_{1/2} - ^2P_{1/2}$



Look for the note! multiply by 2 for D1:



El.	Transition	n_s^*	n_p^*	Theory		Experiment
				μ (a.u.)	Γ (MHz)	Γ (MHz)
H	1s $\rightarrow$ 2p	1.000	2.000	0.745	99.52	99.47
He*	2s $\rightarrow$ 2p	1.689	1.938	2.540	1.64	1.62
Li	2s $\rightarrow$ 2p	1.589	1.959	2.352	5.93	5.92
Na	3s $\rightarrow$ 3p	1.627	2.117	2.445	9.43	10.01
K	4s $\rightarrow$ 4p	1.770	2.235	2.842	5.78	6.09
Rb	5s $\rightarrow$ 5p	1.805	2.293	2.917	5.78	5.56
Cs	6s $\rightarrow$ 6p	1.869	2.362	3.093	4.99	5.18

↑
this is $\langle l\alpha' | r | l\alpha \rangle$

Dan Steck:

includes the signs

	$m_F = -2$	$m_F = -1$	$m_F = 0$	$m_F = 1$	$m_F = 2$
$F' = 3$	$-\sqrt{\frac{1}{6}}$	$-\sqrt{\frac{4}{15}}$	$-\sqrt{\frac{3}{10}}$	$-\sqrt{\frac{4}{15}}$	$-\sqrt{\frac{1}{6}}$
$F' = 2$	$-\sqrt{\frac{1}{6}}$	$-\sqrt{\frac{1}{24}}$	0	$\sqrt{\frac{1}{24}}$	$\sqrt{\frac{1}{6}}$
$F' = 1$		$\sqrt{\frac{1}{40}}$	$\sqrt{\frac{1}{30}}$	$\sqrt{\frac{1}{40}}$	

D2 line,
 $F = 2 \rightarrow F'$
 $\hat{\xi} = \hat{\sigma}$

$$= \frac{\langle F' m_{F'} | r_q | F m_F \rangle}{\langle J=1/2 || r || J=3/2 \rangle}$$

measured: $4.227(5) \text{ a.u.} = 4.227(5) \text{ ea}_0$
 (for D1, $\langle J=1/2 || r || J=1/2 \rangle = 2.992(3) \text{ ea}_0$)
 from excited state lifetimes

$$\frac{\langle F m_F | r_q | F' m_{F'} \rangle}{\langle J \| r \| J' \rangle} = (-1)^{F'-L+m_F} \sqrt{2F+1} \begin{pmatrix} F' & 1 & F \\ m_{F'} & q & -m_F \end{pmatrix} \cdot (-1)^{F'+J+L+1} \sqrt{(2F'+1)(2J+1)} \begin{Bmatrix} J & J' & 1 \\ F' & F & I \end{Bmatrix}$$

and

$$\langle J \| r \| J' \rangle = \langle L \| r \| L' \rangle (-1)^{J'+L+1/2} \sqrt{(2J'+1)(2L+1)} \begin{Bmatrix} L & L' & 1 \\ J' & J & S \end{Bmatrix}$$

$$\frac{\langle J \| r \| J' \rangle}{\langle L \| r \| L' \rangle} = \begin{cases} \sqrt{2} & \text{for D2 line} \\ 1 & \text{for D1 line} \end{cases}$$

Make sure that you include these factors if you use $\langle l' \alpha' | r | l \alpha \rangle$!

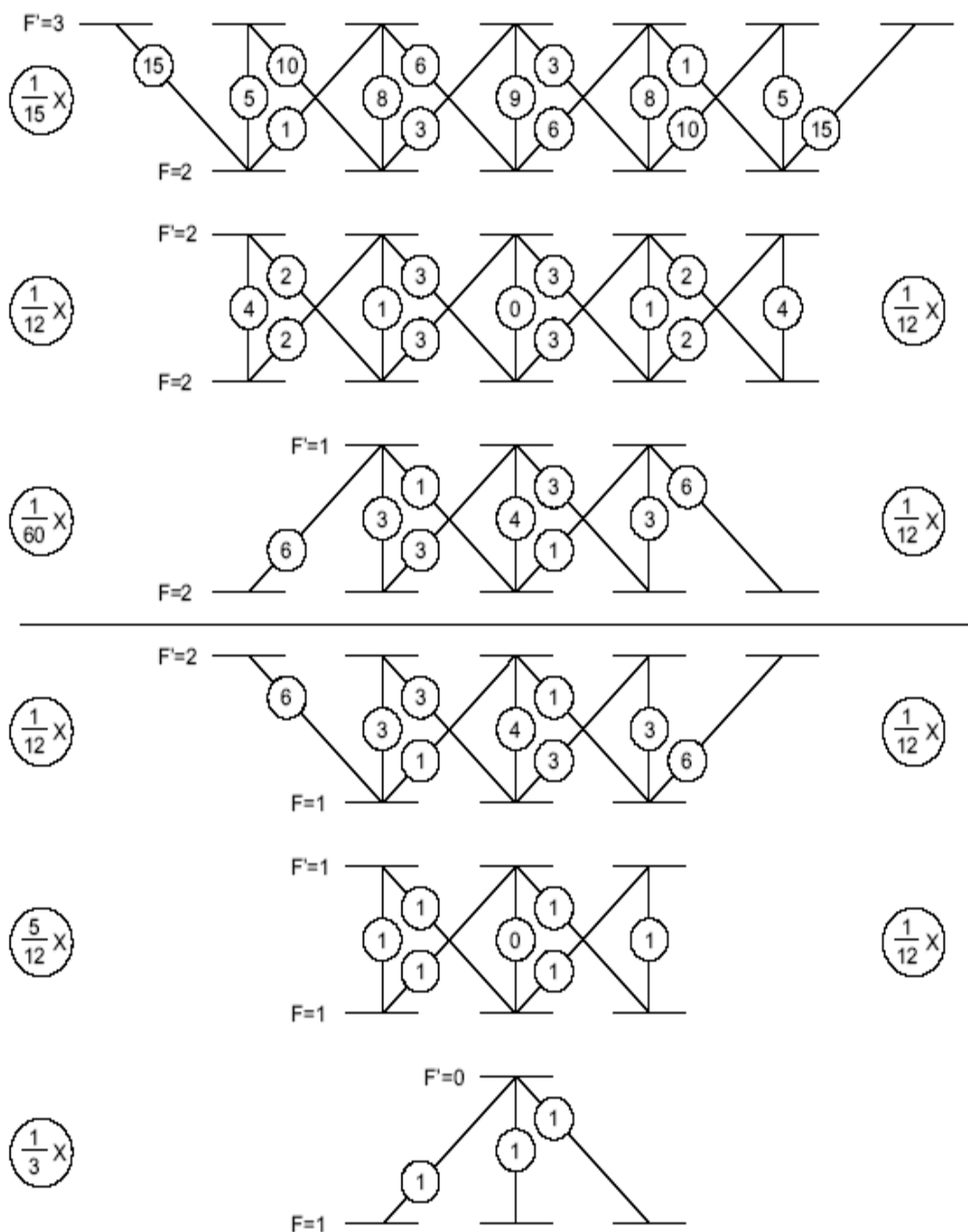
From Eric Cornell / Carl Wieman's groups

No signs!

D2-line (780 nm)

87Rb

D1-line (795 nm)



Take the
square
root

